

Fast and Accurate Non-Linear Predictions of Universes with Deep Learning

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Motivation

- ✓ NN can indeed learn from fast approximation methods the evolution of dark matter particles outperforming analytical methods.

Goal

- ✓ Improvement of this methodology by using high resolution and full N-body simulations for universes with different cosmological parameters.

Types of cosmological simulations

- ✓ **N-body**
 - Brute force method;
 - Computationally expensive.
- ✓ **Linear Theory**
 - [ZA](#), [2LPT](#);
 - Ideal for large-scales, small matter density.
- ✓ **Fast Approximations**
 - Resolves well small-scale issues;
 - [COLA](#), [L-PICOLA](#).
- ✓ **Neural Networks**
 - Outperform Linear Theory and Fast Approximations;
 - Need a dataset for training.

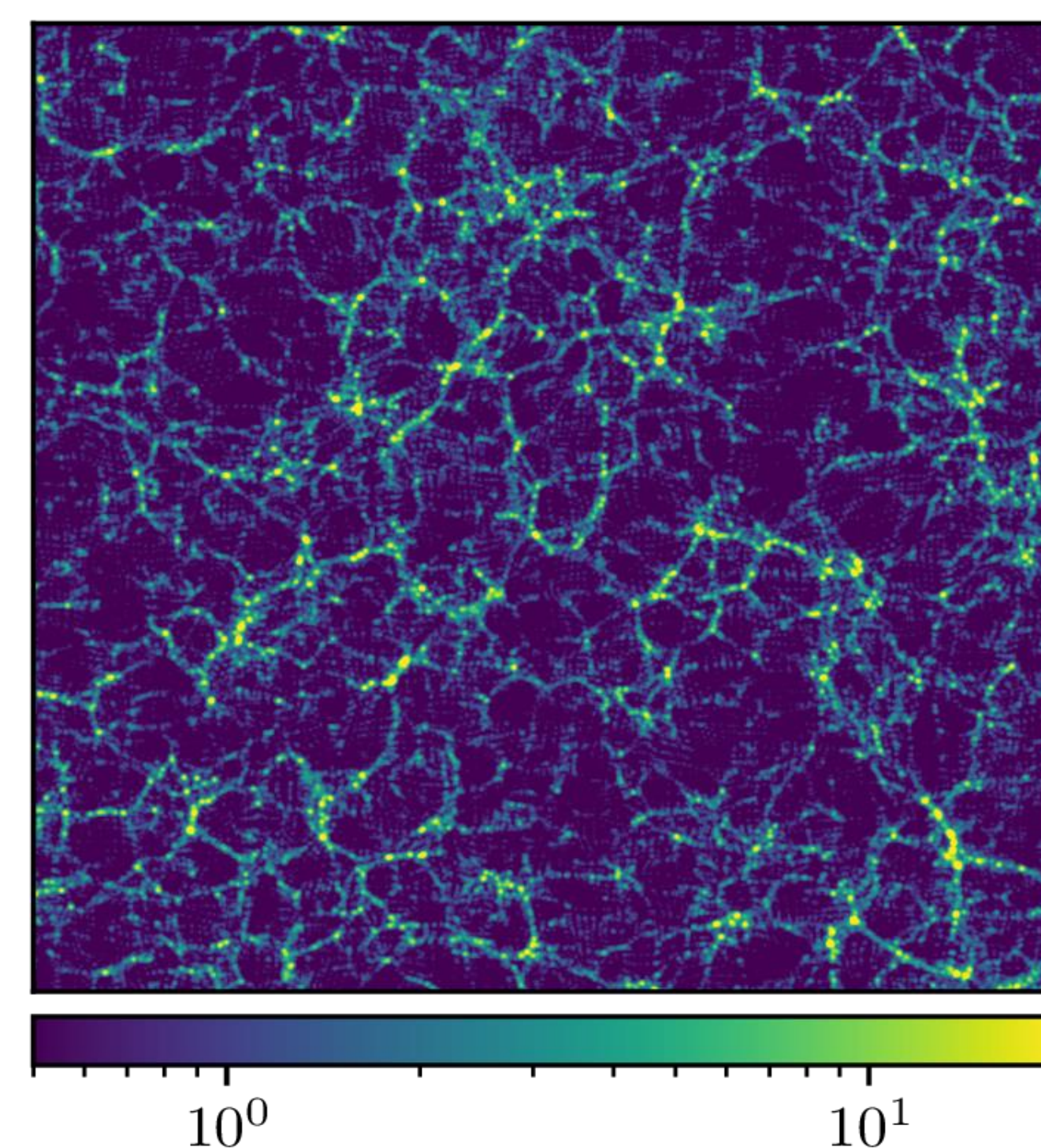
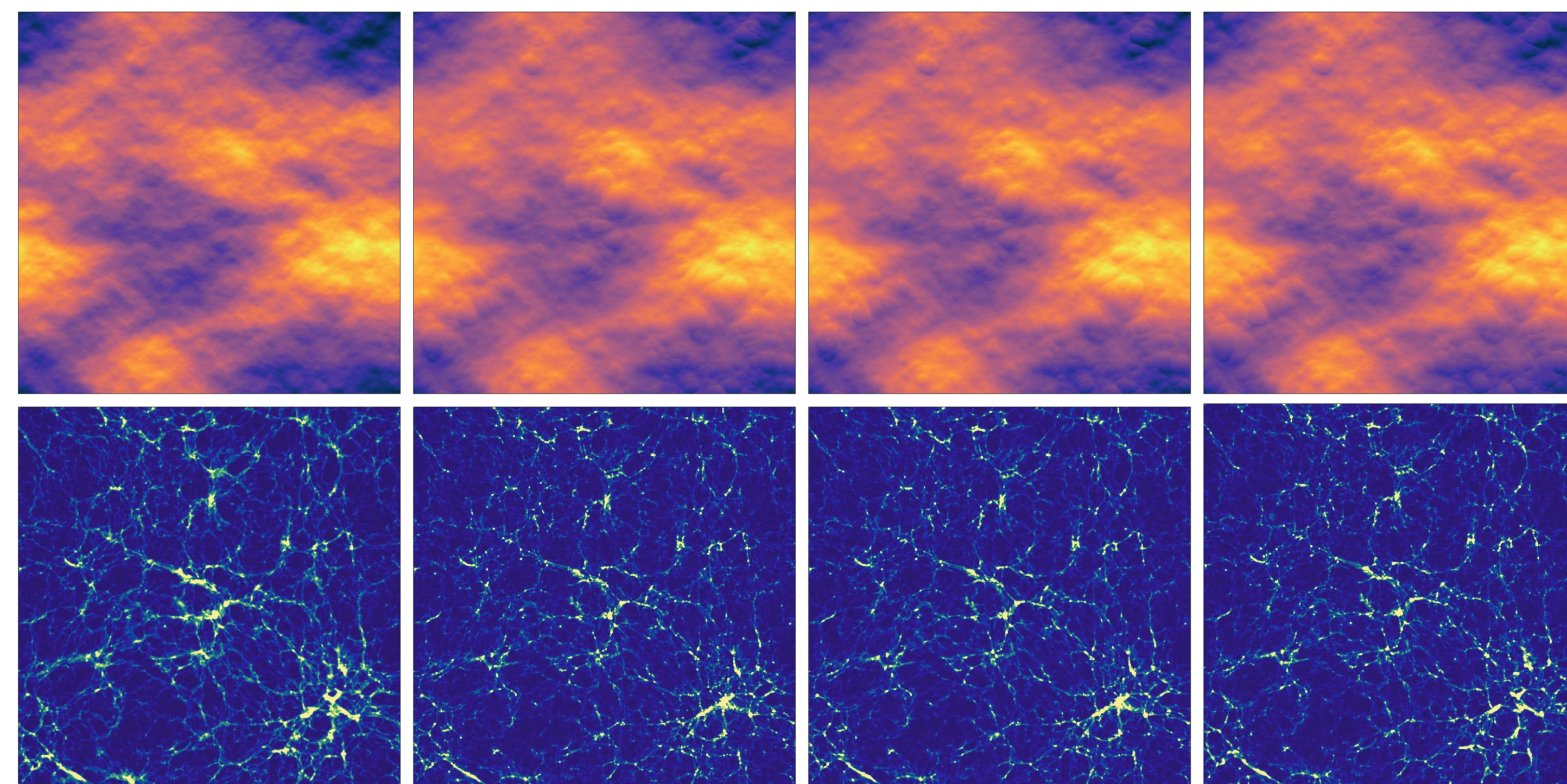


Fig. 1: Slice density δ distribution of dark matter particles in a box of 1 Gpc.

The Model

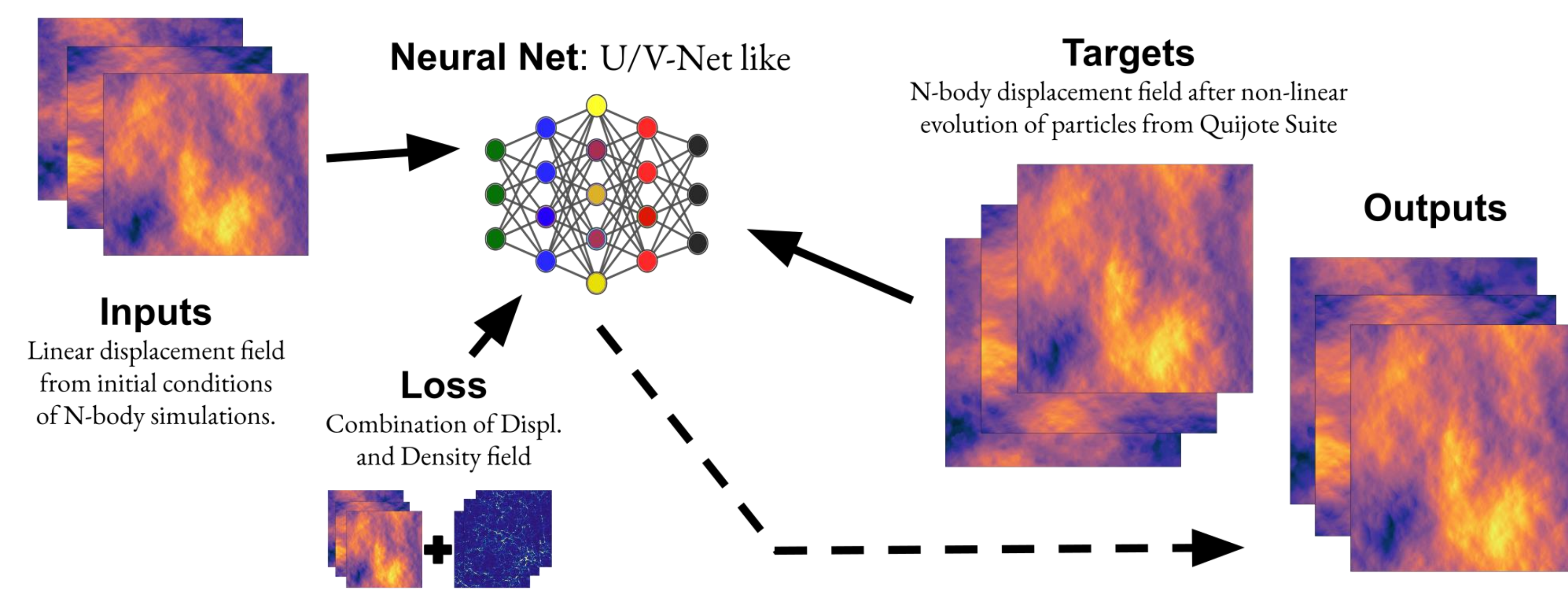


Fig. 2: The first line are slices of displacements of particles and in the second line we show density slices. **Can you spot any difference between these simulations by eye?** From left to right: Linear, N-Body (truth), Fast. Approx., and NN.

Metrics Used

- ✓ Power Spectrum P of Ψ and δ ;
- ✓ Transfer function:

$$T(k) = \sqrt{\frac{P_{\text{pred}}(k)}{P_{\text{true}}(k)}}$$

- ✓ Cross-Correlation:

$$r(k) = \frac{P_{\text{pred} \times \text{true}}(k)}{\sqrt{P_{\text{pred}}(k)P_{\text{true}}(k)}}$$

Results

- ✓ Same set of cosmological parameters: $\{\Omega_m, \Omega_b, \Omega_\Lambda, \sigma_8, h, n_s\}$

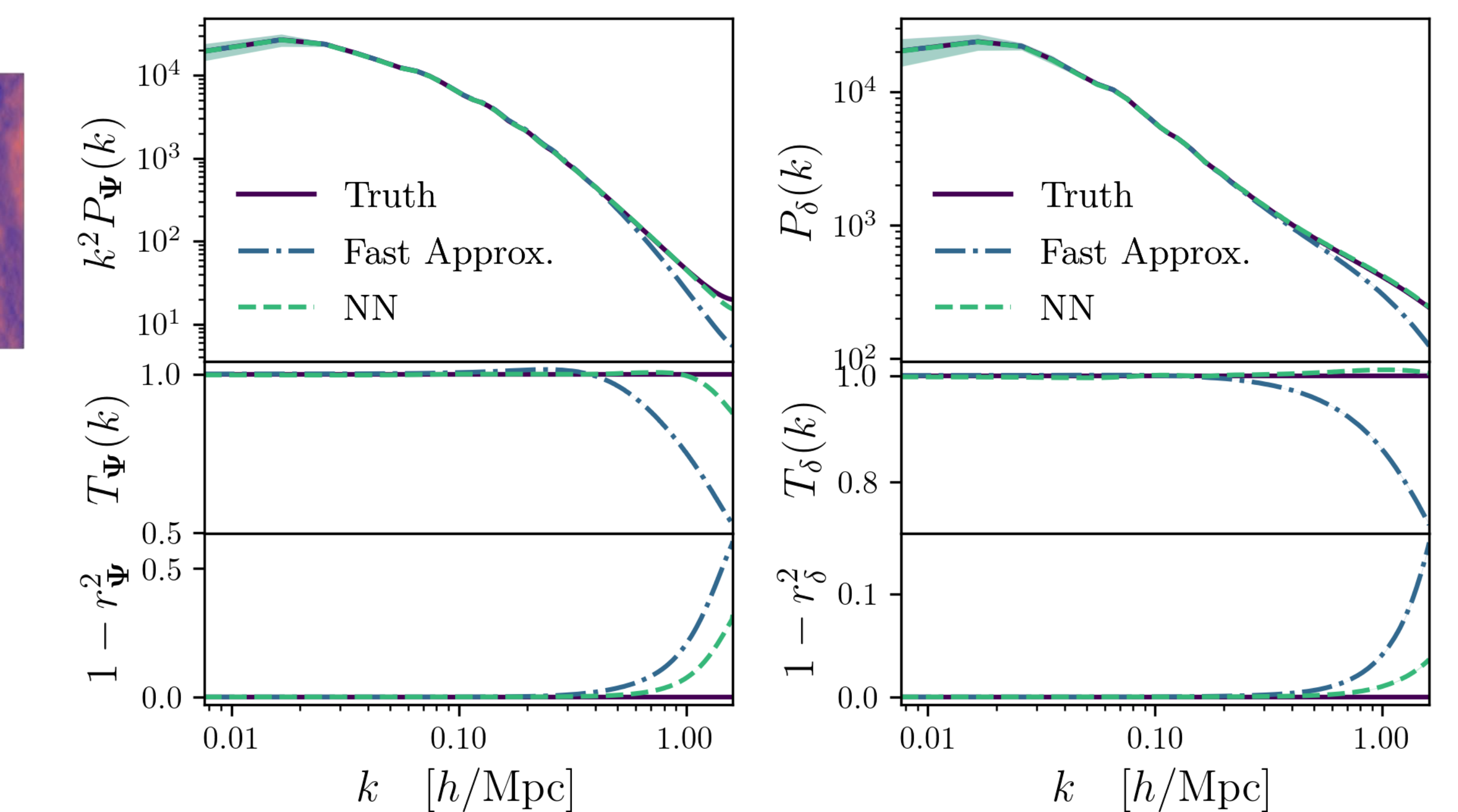


Fig. 3: Metrics for NN and fast approximator model compared with N-body sims.

Other Universes

- ✓ Tested on 2000 types of universes

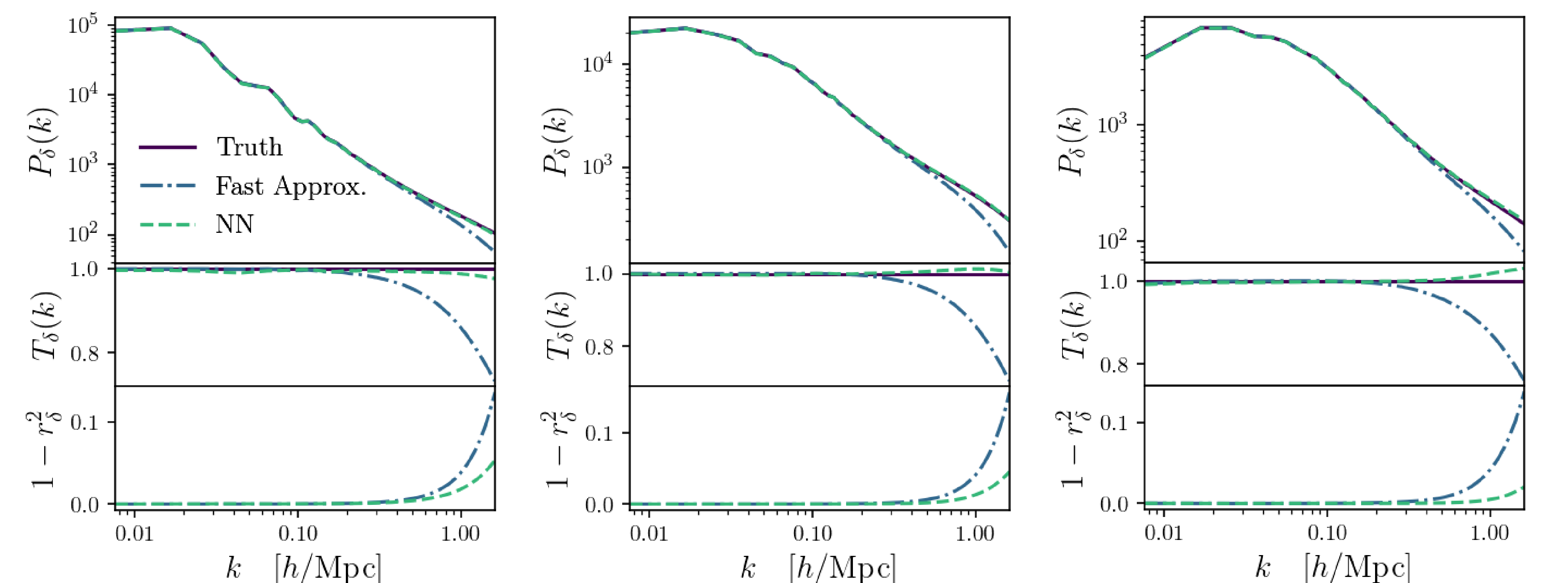


Fig. 4: Accuracy comparison between predictions by the fast approximator (blue dot-dashed) and our NN (green dashed) for universes with different cosmological parameters used in training.